Chapter 2 Waves and Particles

De Broglie wavelength: $\lambda = h/p$, where *p* is the momentum. If it is particle's momentum, then $p = mv = m_0 v/\sqrt{1 - v^2/c^2}$ and $\lambda = h(\sqrt{1 - v^2/c^2})/m_0 v$; else if it is photon's momentum, then p = E/c = hv/c, where $h = 6.626 \times 10^{-34} (J \cdot \text{sec}) = 4.136 \times 10^{-15}$ (eV · sec) is Plank's constant.

Eg. Find the de Broglie wavelengths of (a) a 46g golf ball with a velocity of 30m/sec, (b) an electron with a velocity of $10^7m/\text{sec}$

(Sol.) (a) :: v << c, $\lambda = h/p = h/m_0 v = 6.626 \times 10^{-34}/(46 \times 10^{-3} \times 30) = 4.8 \times 10^{-34}(m)$

(b)
$$v=10^7 m/\text{sec}$$
, $\lambda=h/p=h(\sqrt{1-v^2/c^2})/m_0v=7.3\times10^{-11}(m)$

Quantum energy of one photon: $E=hv=hc/\lambda=1.24\times10^{-6}/\lambda=pc$. Static mass of one photon: $m_0=hv/c^2$.

2-1 Photoelectric Effect *E=hv=K*_{max}+*hv*₀



Photoelectric Effect: $E=hv=K_{max}+hv_0=Maximal$ electron's energy + Work function Eg. An UV light of wavelength 350nm and intensity $1W/m^2$ is incident at the potassium surface. (a) Find the maximum kinetic energy of photoelectrons. (b) If 0.5 percentages of the incident photons produce photoelectrons, how many photoelectrons/sec are emitted if potassium surface has an area of $1cm^2$? (Sol.) (a) Work function of $K=2.2eV=hv_0$. $E=1.24\times10^{-6}/350\times10^{-9}=3.5eV$, $K_{max}=E-hv_0=3.5-2.2=1.3eV$

(b) $3.5eV=3.5\times1.602\times10^{-19}=5.68\times10^{-19}J$, $n_p=\frac{1\times10^{-4}}{5.68\times10^{-19}}=1.76\times10^{14}$ photons/sec $1.76\times10^{14}\times0.5\%=8.8\times10^{11}$ photoelectrons/sec Eg. Blue light of wavelength 300nm and intensity $0.5W/m^2$ is directed at a material with work function of 2.7eV. (1) Find the maximum kinetic energy of photoelectrons. (2) If the quantum efficiency is 1%, how many photoelectrons are emitted per second if the semiconductor surface has an area of $2cm^2$? [台大電 研]



An x-ray tube. The higher the accelerating voltage V, the faster the electrons and the shorter the wavelengths of the x-rays.

The shortest wavelength of X-ray due to electron bombard: $\lambda_{\min} = 1.24 \times 10^{-6}/V$, where V is the accelerating voltage.

Eg. Find the shortest wavelength present in the radiation from an X-ray machine whose accelerating potential is 50000V.

(Sol.) $\lambda_{\min}=1.24\times10^{-6}/50000=2.5\times10^{-11}m=0.025nm$

2-2 Bragg's Reflection $2d\sin\theta = n\lambda$ and Compton Effect $\Delta \lambda = h(1-\cos\varphi)/m_0c$ Bragg's Reflection: When $2d\sin\theta = n\lambda$ occurs, it forms a constructive interference.



Eg. The distance between adjacent atomic planes in calcite is 0.3*nm*. What is the smallest angle between these planes and an incident beam of 30*pm X*-rays at which scattered *X*-rays can be detected?

(Sol.) $d=0.3\times10^{-9}$, $\lambda=30\times10^{-12}$, $2d\sin\theta=n\lambda=\lambda\Rightarrow\sin\theta=0.05$, $\theta=2.86^{\circ}$

Compton Effect: $\triangle \lambda = h(1-\cos\varphi)/m_0c$ $\begin{cases} hv/c+0 = hv'\cos\varphi/c + p\cos\varphi \\ 0 = hv'\sin\varphi/c - p\sin\varphi \end{cases} \Rightarrow (pc)^2 = (hv)^2 - 2hv \cdot hv'\cos\varphi + (hv')^2 \dots (1)$ And $E = K + m_0c^2 = (hv-hv') + m_0c^2 \Rightarrow (pc)^2 = (hv)^2 - 2hv \cdot hv' + (hv')^2 + 2m_0c^2(hv-hv') \dots (2)$ (1), (2) $\Rightarrow (m_0c/h)(v/c-v'/c) = (v/c)(v'/c)(1-\cos\varphi) \Rightarrow \lambda' - \lambda = \Delta = h(1-\cos\varphi)/m_0c$



Compton wavelength: $\lambda_c = h/m_0 c$, and the Compton wavelength of the electron is $\lambda_{c.electron} = 2.426 \times 10^{-12} m$

Eg. A beam of X-rays is scattered by a target. At 45° from the beam direction the scattered X-rays have a wavelength of 2.2pm. What is the wavelength of the X-rays in the direct beam?

(Sol.) $\varphi = 45^{\circ}, \lambda' = 2.2pm, h/m_0c = 2.426 \times 10^{-12}, \lambda' - \lambda = h(1 - \cos\varphi)/m_0c \Longrightarrow \lambda = 1.54pm$

2-3 Photon, Gravity, and Black Holes



Eg. Find the change in frequency of a photon of red light whose original frequency is $7.3 \times 10^{14} Hz$ when it fall through 22.5m.

(Sol.) H=22.5, $\triangle v=gHv/c^2=1.8Hz$



Gravitational red shift: The frequency of a photon emitted from the surface of a star decreases as it moves away from the star. And its total energy becomes $E=hv-GMm/R=hv-GM(hv/c^2)/R=hv'$

$$\Rightarrow \Delta v/v = (v - v')/v = = (1 - v'/v) = \frac{GM}{c^2 R}$$

- (1) For most stars, M/R is small, $\Delta v/v$ is not apparent.
- (2) For white dwarf stars, M/R is large, $\Delta v/v$ is measurable.
- (3) For black holes, if $\frac{GM}{c^2R} \ge 0.5$, we can obtain Schwarzschild radius $R_s=2GM/c^2$,

and prove that no photon can leave the black holes.